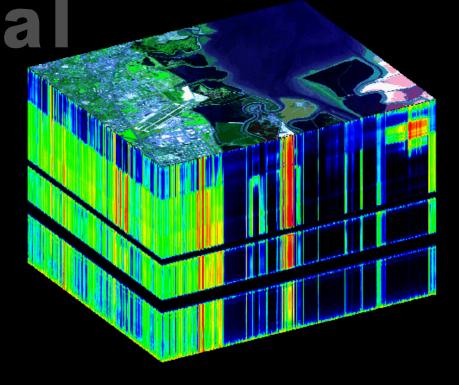


Hyperspectra Challenge Problem

Adapted in part from charts by R. Taylor, NASA GSFC

With helpful input from Pankaj Topiwala USC/ISI

Specific Hyperspectral Challenge Problem: LANL

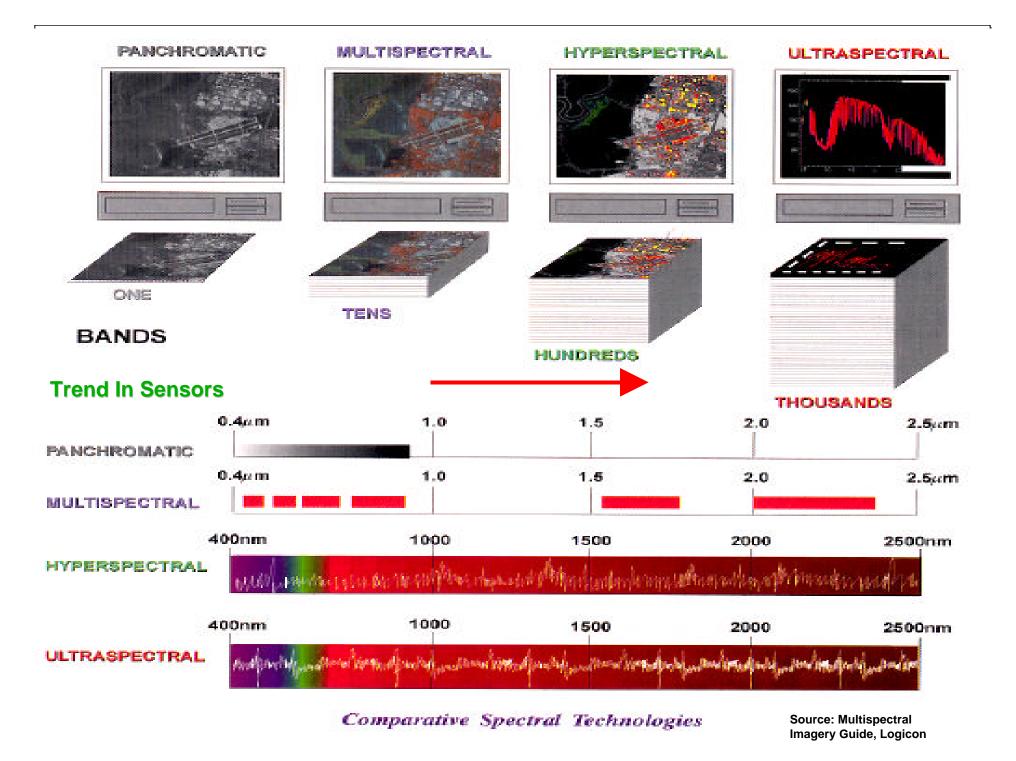


Jeffrey Bloch
Los Alamos National Laboratory
jbloch@lanl.gov

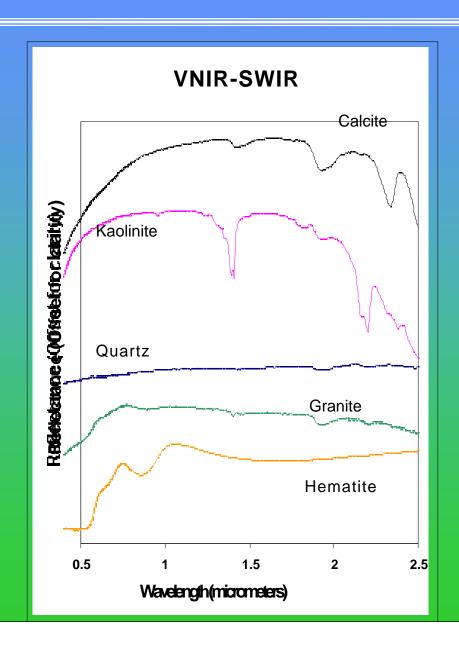
Objective

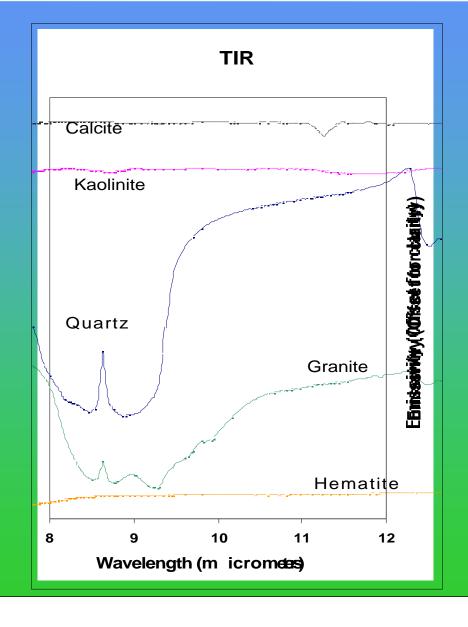
- Present the motivation and a long-term vision for highresolution spectral imaging
- Present a detailed challenge problem for onboard, hyperspectral product generation using reconfigurable computing





Spectral Signatures





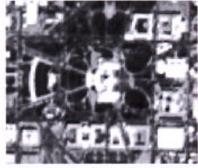
Orbview Pan - 1 meter GSD



Orbview Color - 4 meter GSD



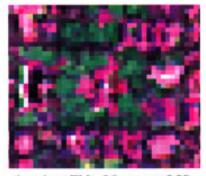
IRS-1C Pan - 6 meter GSD



SPOT Pan - 10 meter GSD



SPOT XS - 20 meter GSD



Landsat TM - 30 meter GSD

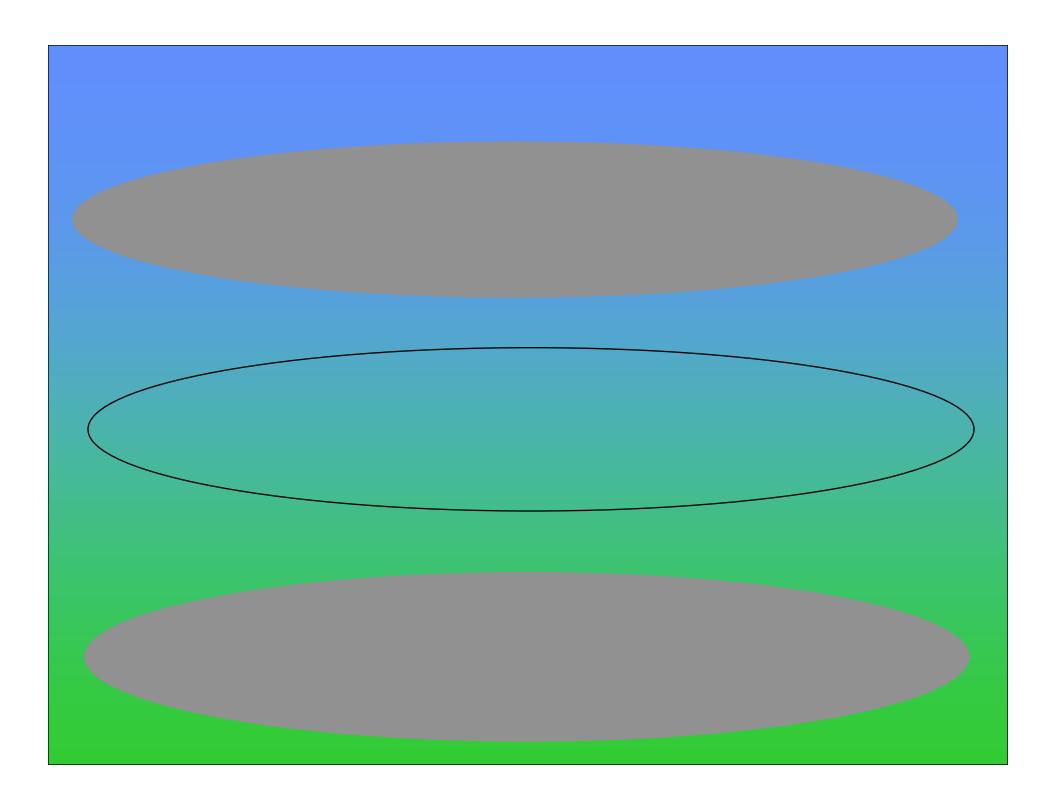
Images Provided and Reproduced by Permission of OrbView, Space Imaging EOSAT, and ⊕ CNES/SPOT Image.

Source: Multispectral Imagery Guide, Logicon

Spatial Vs. Spectral Resolution

- Sensors can trade spatial vs spectral resolution
- Planned sensors have high spatial AND spectral resolution
- Exploitation is processor/storage intensive

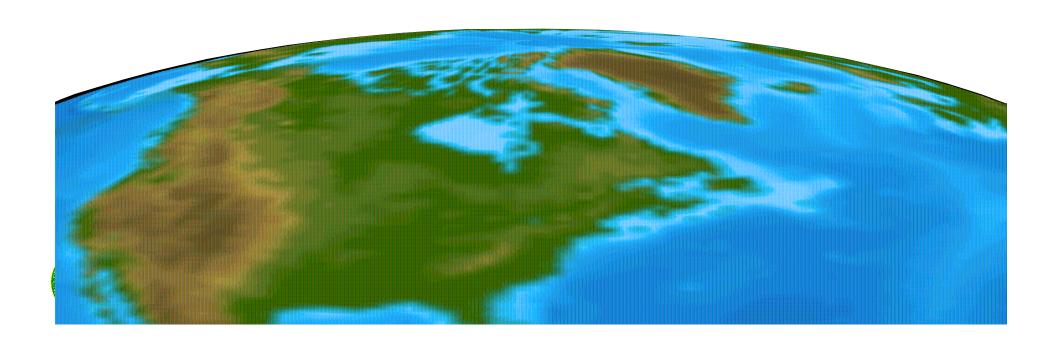




Near-Term Vision: Hyperspectral Imaging System

DATABASE-IN-THE-SKY

- •Users Request Hypercube Slices
- Direct Delivery of Derived Products
- •Large Area Coverage per Orbit
- •Full Spectrum Coverage (VNIR TIR)
- Pointable for Revisit / Calibration



Near-Term Vision: Hyperspectral Imaging System

SYSTEM SPECIFICATIONS

•Orbit period: 90 min

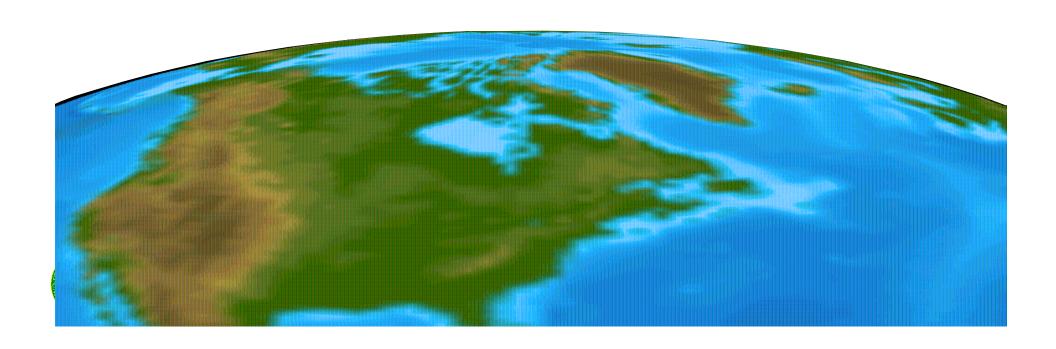
•Swath width: 5 km - 500 km

•Spatial Resolution: 2.5 m - 250 m

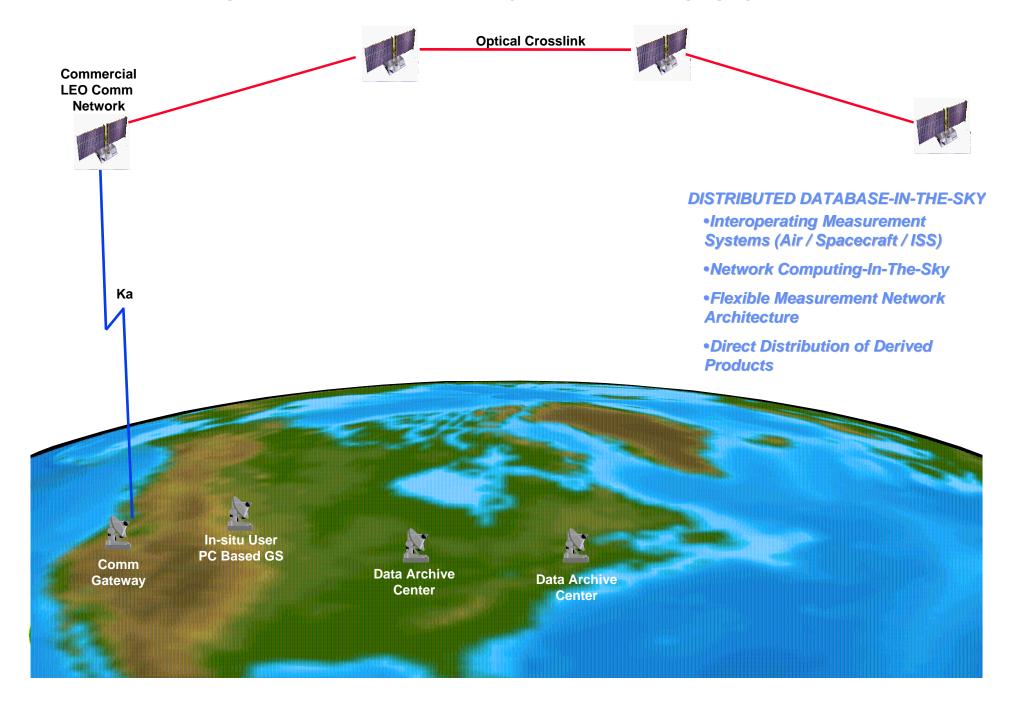
•Spectral Resolution: 200 bands

•Dynamic Range: 12 bpp

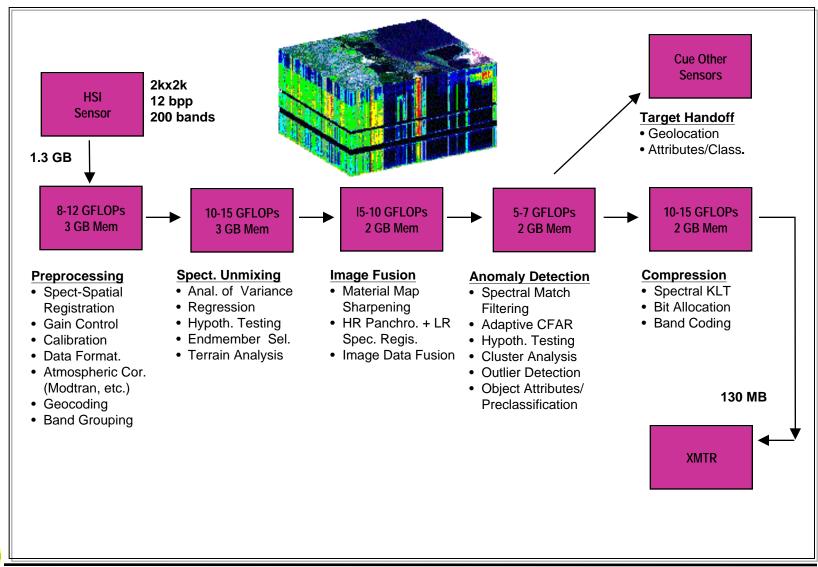
•Data Rate: up to 2 GB/s



Long-Term Vision: Distributed Hyperspectral Imaging Systems



Generic Characteristics of a Hyperspectral Image Processing Pipeline





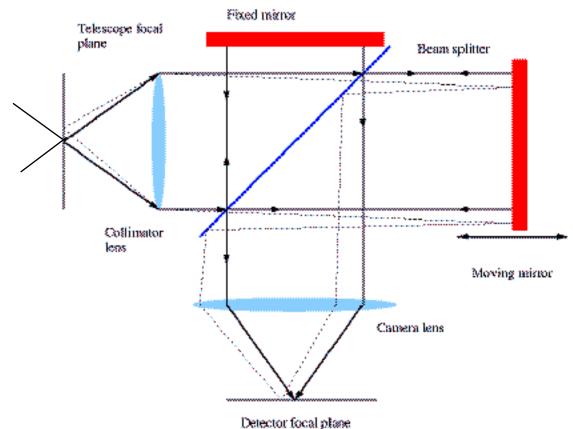
The Challenge Problem: Real-Time Continuous Hyperspectral Surveillance

- Assumptions:
 - Fourier Transform Imaging Instrument
 - Datacube axes: X, Y, and F.T. of optical spectrum
 - New image cube collected every 3 seconds
- Real Time Processing Requirements:
 - Recover pixel spectra
 - Search for multiple spectral features of interest
 - (32 Maximum using matched filters)
 - Final output: Broadband and spectral feature images
 - Extra Credit: Image slice motion correction
 - Extra Credit: Pixel Clustering in real time



Challenge Problem Data Source: Fourier Transform Imaging Spectrometer

Imaging Michelson Interferometer



Instrument Data Characteristics:

- Spatial Resolution: 128x128 pixels
- 14 Bits of data per pixel
- Mirror Positions: 4000+
- Number of mirror positions not necessarily a power of 2
- Mirror Scan Time over all locations: 3 seconds
- Each spatial image is collected before the mirror is moved again.
- Image Data cube of contains an *Interferogram* of the optical spectrum at each spatial pixel.
- The FOV of the instrument *may* move slightly between image collections. Such translations need to be removed, preferably at the sub-pixel level





Algorithm Overview

All of the following must be performed in less than 3 seconds:

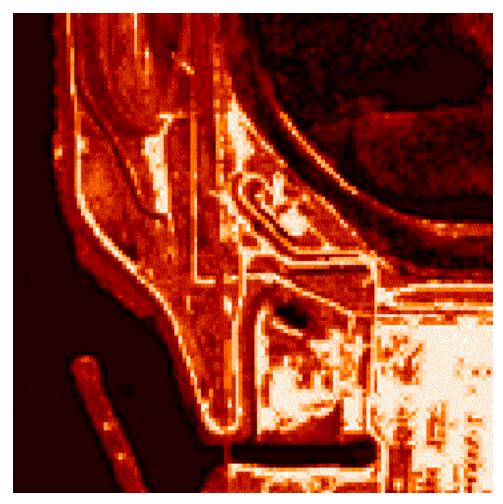
- Step 1A: Extra Credit: Register each image slice with the next
- Step 1: Perform a FT on each pixel's Interferogram
 - -> 1634 4096 pt FTs
- Step 2: Combine and average symmetric portions of transforms
 - Actual number of valid wavenumber bins is 251
- Step 3: Apply wavelength and pixel dependent gains and offsets
 - Calibration table supplied as part of the challenge problem
- Step 4: Sum the datacube in the wavenumber dimension and produce a broadband image
- Step 5: Compute the Covariance Matrix of the image background
- Step 6: Use the covariance matrix on 32 library spectral templates to produce 32 matched filters.
- Step 7: Produce 32 images by performing a spectral dot product of the matched filters on the datacube.
- Step 8: Extra Credit: Perform a "k-means" cluster assignment algorithm on the spectral datacube, classifying each pixel, and reporting the mean class vector for each cluster as well as an image with the assignments.

NOTE: Steps 5+6 can be bypassed in a first implementation.



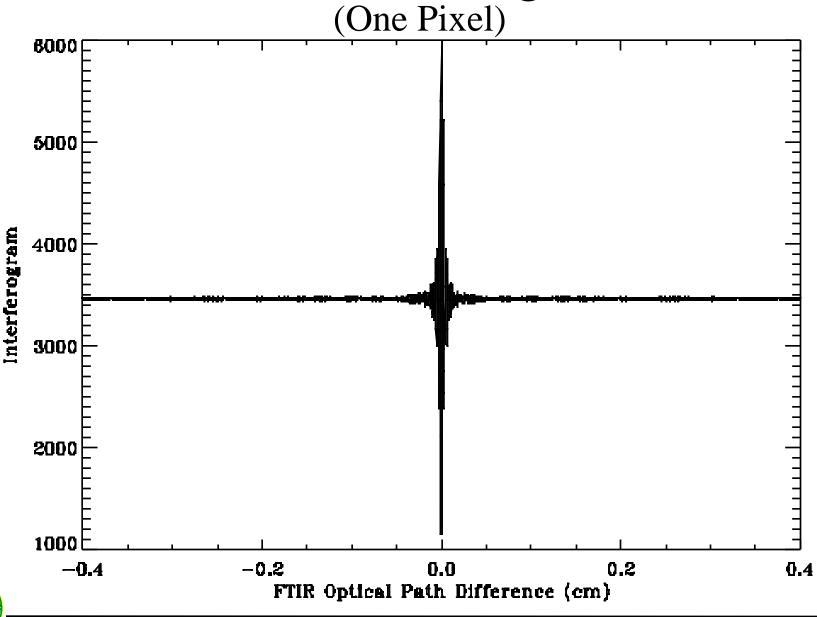
Broad Band Image

(All Spectral Channels Summed Together)



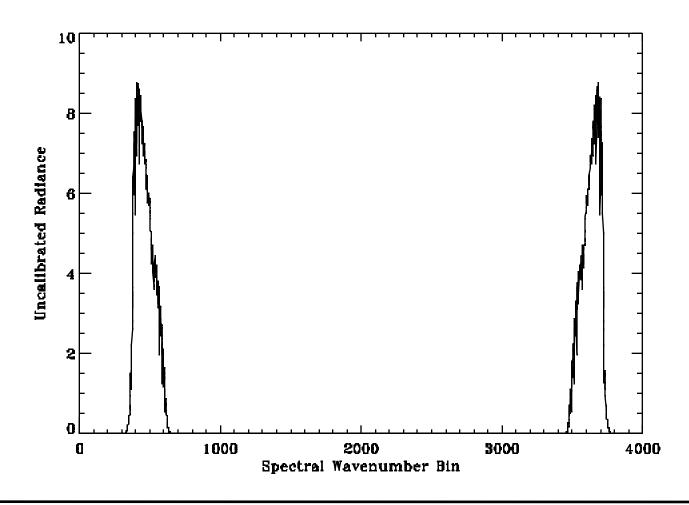


Raw Interferogram





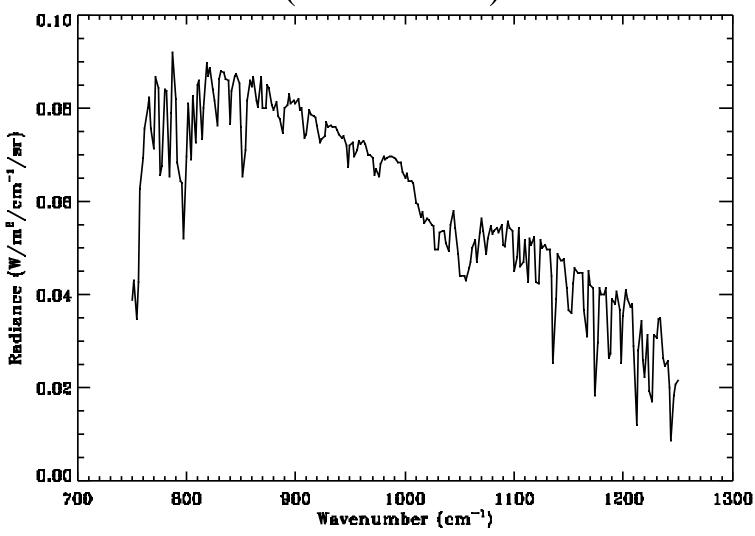
Fourier Transform of Interferogram





Final Calibrated Spectrum

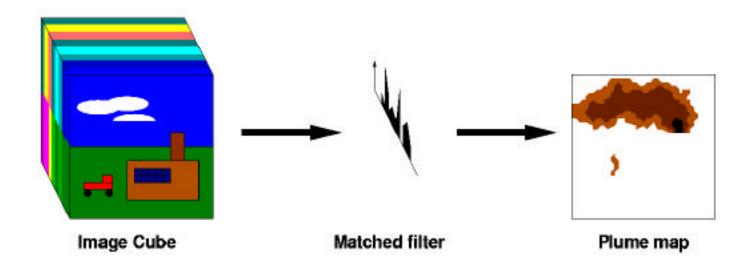
(for one Pixel)





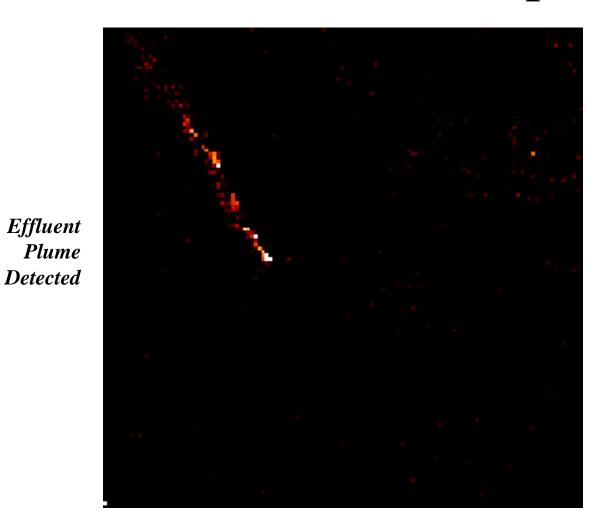
Spectral matched filter

- $\mathbf{R} = \text{image cube}$
- b = spectral signature of plume
- $\mathbf{K} = \mathbf{R}\mathbf{R}^T = \mathbf{K}$ covariance matrix; describes statistics of background clutter
- $\mathbf{q} = \mathbf{b} =$ ordinary matched filter
- $\mathbf{q} = \mathbf{K}^{-1} \mathbf{b} = \text{adaptive matched filter}$
- $\mathbf{q}^T \mathbf{R} = \text{plume map}$



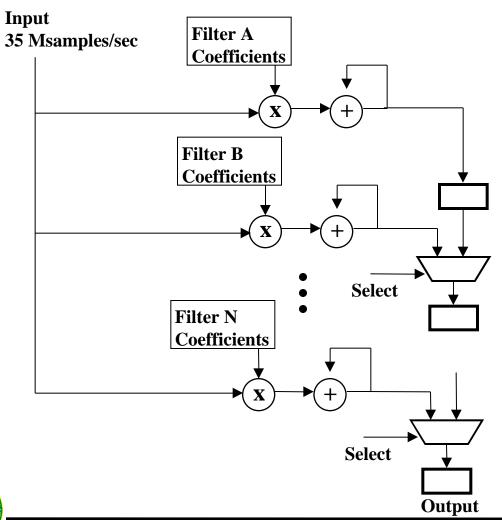


Matched Filter Output





Spectral Matched Filter on LANL Reconfigurable Computing Board



- •Current Hardware implementation 17x faster than 400 MHz Intel PII for 12 filters.
- •Fully populated Hardware implementation 45x faster than 400 MHz Intel PII for 32 filters.
- More processing capacity available for additional operations (classification, thresholding, etc.) in parallel.
- Multiple boards can be ganged together, increasing processing while maintaining throughput.



Challenge Problem on Web

Details of this hyperspectral ACS challenge problem, including the sample raw dataset, calibration constants, sample output, and algorithm descriptions will be available on the web in approximately 3 weeks.

Please send e-mail to:

jbloch@lanl.gov

and you will be notified when the web site is available.

